Topics 19-20

Spectroscopy

Fall 2016

SPECTROSCOPY: short wavelength regions ESCA (photoelectron) and UV

handout

approach for spectroscopy

- material not straight from text chapter
- must FOLLOW videos, "lectures", handout and worksheet
- (WA) HW is from SAMPLE FINAL QUESTIONS on spectroscopy
- for discussion group (M-Tu 28-29 November): inquiry exercise Galen Gorski (UCSC graduate student EarthSci, ISEE)



Spectroscopy Handout

1. Know what type of spectroscopic technique is used to probe each of the excitations described in the chart below:

λ (nm)	υ (sec ⁻¹)	radiation	technique	molecular excitation
.3	10 ¹⁸	x-rays	ESCA	excitation of inner shells breaking of bonds (x-ray damage)
30	10 ¹⁶	far uv	vacuum UV	excitation of σ electrons
300	10 ¹⁵	near uv	UV-VIS	excitation of π and
400-70	00 4-8 × 10 ¹⁴	visible	00-015	non-bonding (n) electrons
3000	10 ¹³	infra-red	IR	vibrational excitations (IR)
			microwave	rotations of molecules and
3 × 10	⁶ 10 ¹¹	microwave	ESR	flipping unpaired electron spins in external magnetic field
3 × 10	9 10 ⁸	radiowave	NMR (MRI)	flipping of nuclear spins in an external magnetic field

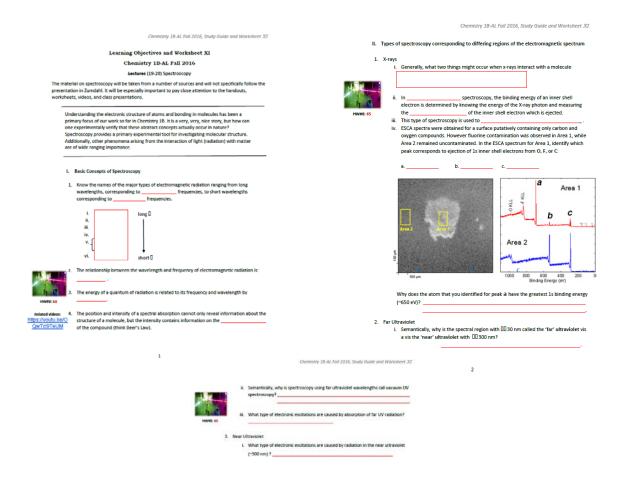
- 2. How are NMR and IR used to elucidate molecular structure?
- 3. What type of molecules have electronic absorptions in the visible region of the
- 4. Know the meaning of:

 - a. Fluorescence
 b. Phosphorescence
 - c. Radiationless decay d. Chemiluminescence

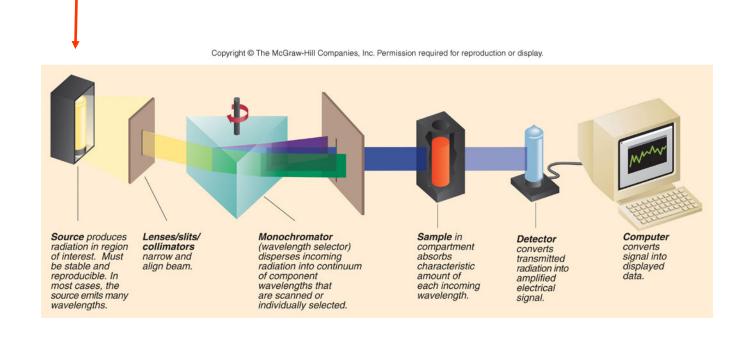
Page 1 of 4

clicker questions worksheet 11, sections I-II.3





spectroscopic principles (Chem 1M/1N exps. #6, #9 and #11)



spectroscopic excitations ($E = hv = hc/\lambda$; $\lambda v = c$)

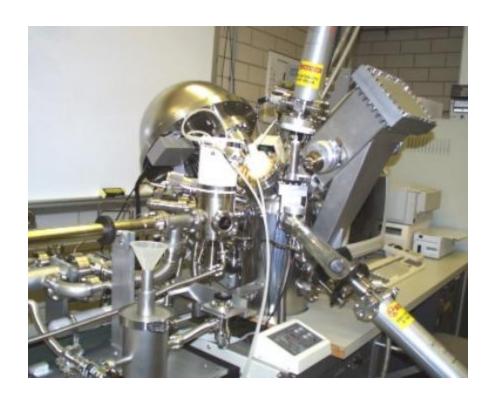
λ (nm)	υ (sec ⁻¹)	radiation	technique	molecular excitation

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400-700	4-8 × 10 ¹⁴	visible	0V-VI3	non-bonding (n) electrons
3000	10 ¹³	infra-red	IR	vibrational excitations (IR)
3 × 10 ⁶	10 ¹¹	microwave	microwave ESR	rotations of molecules and flipping unpaired electron spins in external magnetic field
3 × 10 ⁹	108	radiowave	NMR (MRI)	flipping of nuclear spins in an external magnetic field

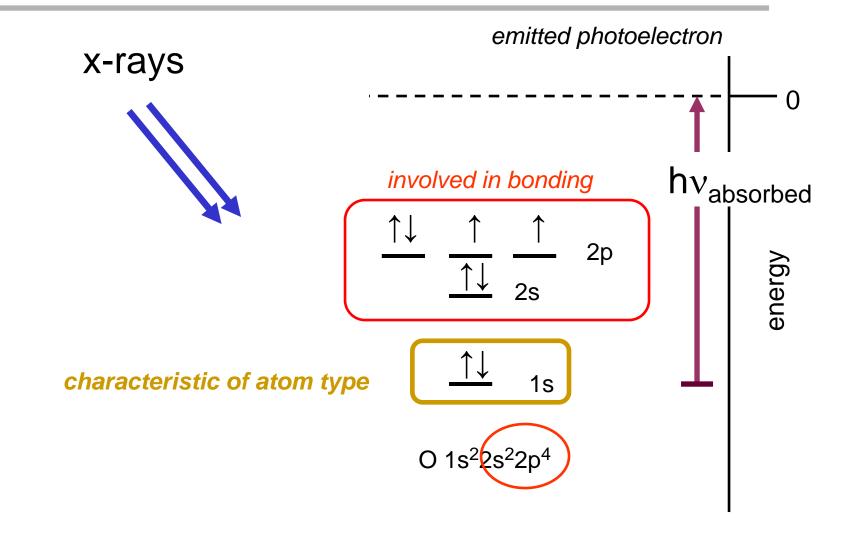
spectroscopic excitations: ESCA

λ (nm)	υ (sec -1)	radiation	technique	molecular excitation
.3	10 ¹⁸	x-rays	ESCA	excitation of inner shells breaking of bonds (x-ray damage)

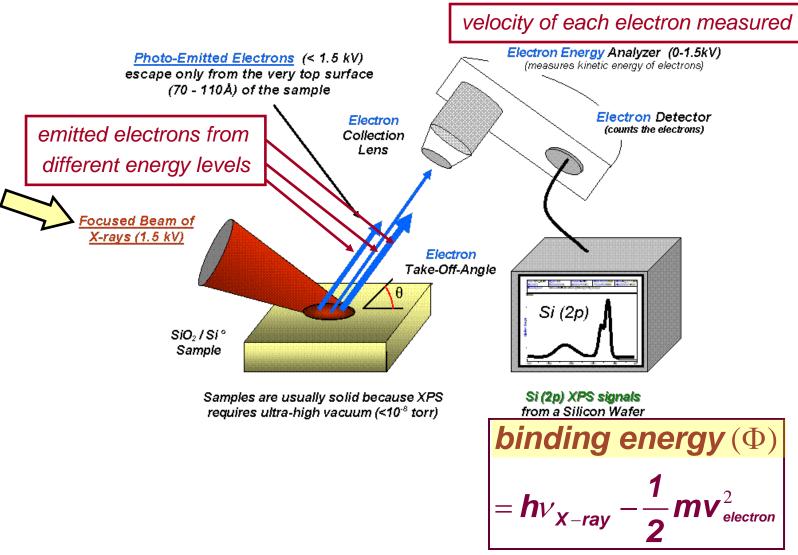
Electron Spectroscopy for Chemical Analysis



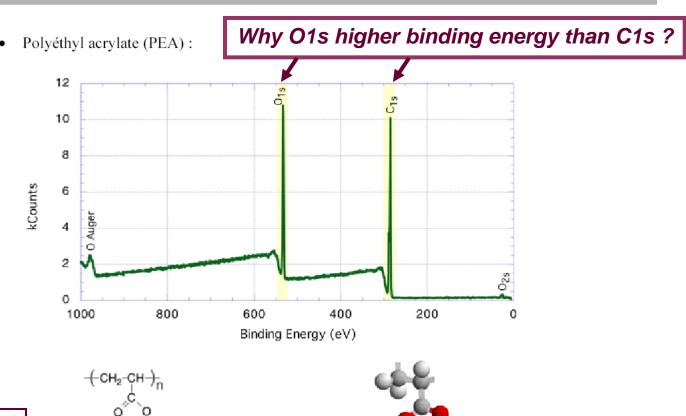
ESCA- photoelectric effect for inner shells

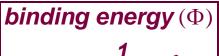


ESCA (and photoelectron effect)



ESCA (binding energy is like work function for inner electrons)





$$= h v_{X-ray} - \frac{1}{2} m v_{electron}^2$$



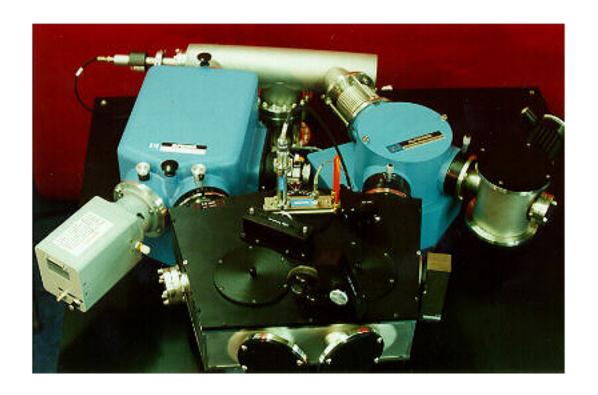
Atome	Composition (%)
Carbone	72,7
Oxygène	27,3

spectroscopic excitations: ESCA

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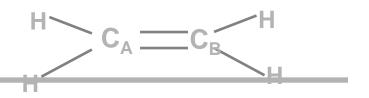
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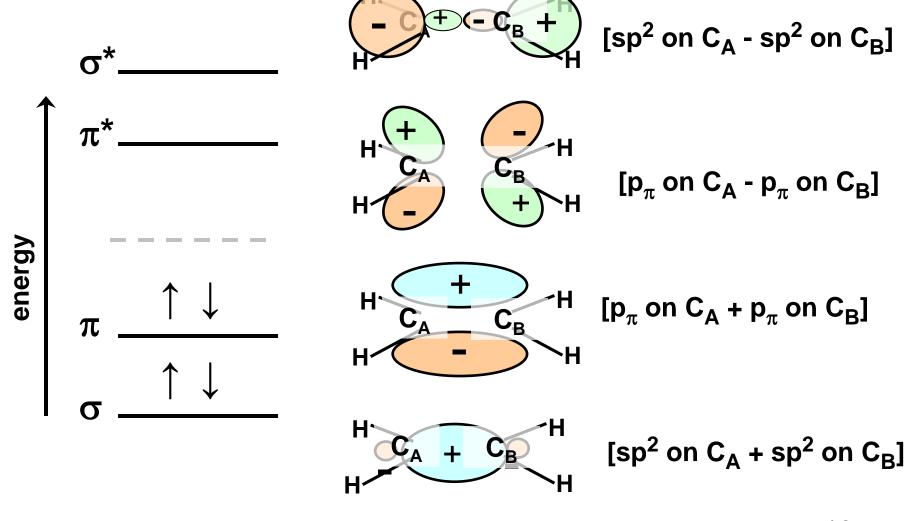
Mc Pherson Vacuum UV Spectrophotometer System



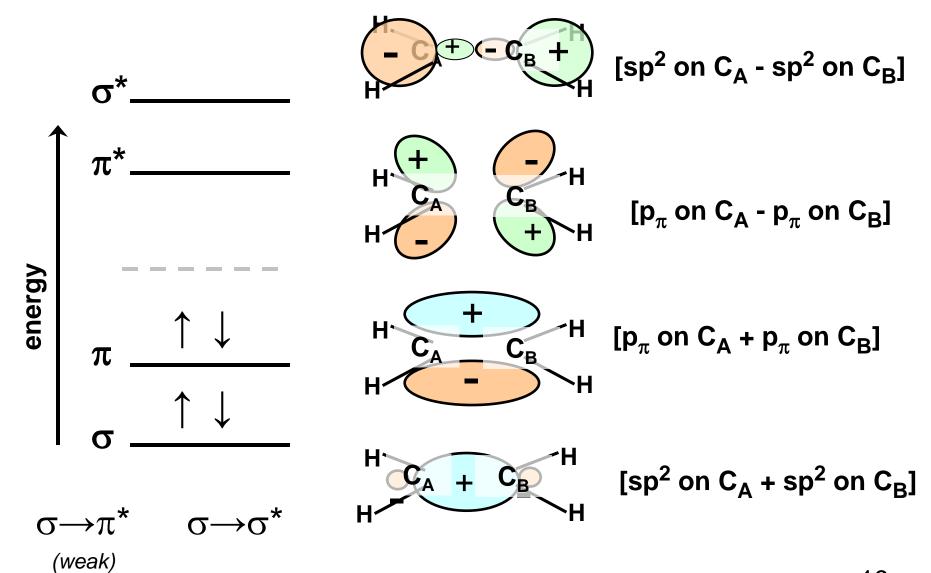
- destructive interference leads to antibonding orbitals which are not usually occupied in the ground state of molecules but which may become occupied upon excitation of electrons by light
- types of antibonding orbitals:

CH₄:
$$\sigma^* = sp^3$$
 on C – 1s on H
C₂H₄: $\sigma^* = sp^2$ on C_A – sp^2 on C_B
 $\pi^* = p_{\pi}$ on C_A – p_{π} on C_B



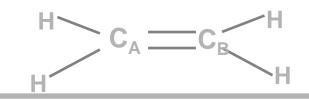


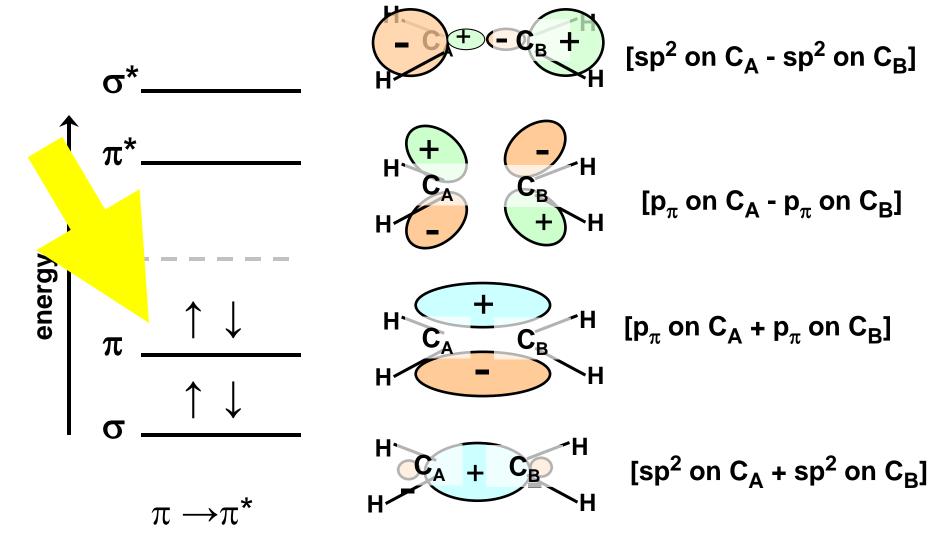
$\square \rightarrow \square^*$ and $\square \rightarrow \pi^*$ excitations are in the far UV regrion

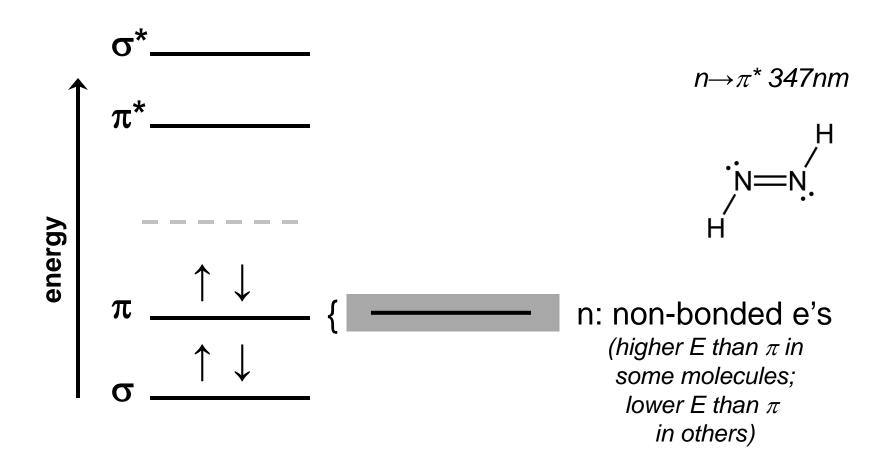


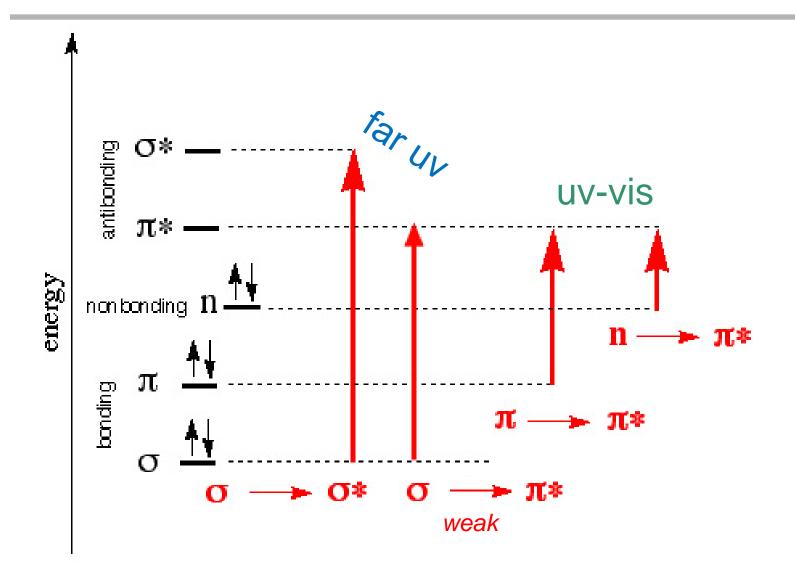
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400-700	4-8 × 10 ¹⁴	visible	07-713	non-bonding (n) electrons









spectroscopic excitations (UV-VIS)

λ (nm)	υ (sec ⁻¹)	radiation	technique	molecular excitation
.3	10 ¹⁸	x-rays	ESCA	excitation of inner shells breaking of bonds (x-ray damage)
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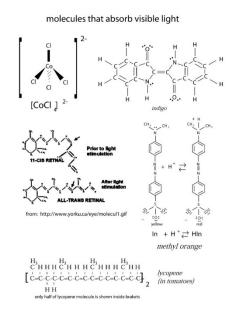
UV-VIS spectrometers (Chem 1M/1N exps. #9 and #11)







why do objects appear colored ?? (previously) what molecules absorb visible light ??















this video on short wavelength spectroscopic techniques was extracted from a full lecture length movie produced for a previous UCSC CHEM1B

flip class





this week
WebAssign HW9
spectroscopy problems from
sample final (conceptual)
due Tuesday November 29

next week
WebAssign HW10
kinetics calculations
(last WebAssign)
due Sunday December 4



why do objects appear colored ??

low energy electronic absorptions in the 'visible' region of electromagnetic spectrum result in the reflection (transmission) of wavelengths of the complementary color.



Highest Occupied Molecular Orbital → Lowest Unoccupied Molecular Orbital

absorbs photon of energy h□ =E_{LUMO}-E_{HOMO}

molecules that absorb visible light

$Br_2(g), I_2(g), NO_2$

[CoCl \rfloor ²⁻ indigo closely spaced HOMO and LUMO due to d-orbital m.o.s or open shells (unpaired e's)

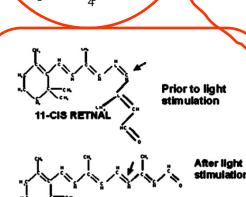
transition metal complex ions octahedral complex $t_{2g} \rightarrow e_g$

(lone-pair) $n \rightarrow \pi^*$

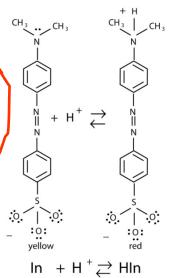
 $\pi \to \pi^*$ in molecules

with conjugated pi-systems

rhodopsin, the molecule most important to seeing color



from: http://www.yorku.ca/eye/molecul1.gif



methyl orange

half of lycopene molecule is shown inside brakets

spectroscopic excitations

λ (nm)	υ (sec -1)	radiation	technique	molecular excitation
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400-700	4-8 × 10 ¹⁴	visible	00-013	non-bonding (n) electrons
3000	10 ¹³	infra-red	IR	vibrational excitations (IR)

diatomic molecules (vibrating and rotating)





http://www.tau.ac.il/~phchlab/experiments_new/LIF/theory.html

wave number

new 'unit'

$$\tilde{v} = \frac{1}{\lambda_{cm}} = \frac{v}{c} \quad [\tilde{v}] = cm^{-1}$$

$$E = hc\tilde{v}$$
 $c = 2.99792 \times 10^{10} cm sec^{-1}$

larger $\tilde{v} \Rightarrow$ higher energy photon larger $\tilde{v} \Rightarrow$ higher energy vibration

vibrational frequency ($ilde{ u}$) depends on

- mass of atoms (lighter \Longrightarrow higher $\tilde{\nu}$)
- strength of bond (tighter 'spring' \Longrightarrow higher $ilde{
 u}$)

b.o.

wave number

larger $\tilde{v} \Rightarrow$ higher energy photon

vibrational frequency ($ilde{ u}$) depends on

- mass of atoms (lighter \square higher $\widetilde{\nu}$)
- ullet strength of bond (tighter 'spring' llet higher $ilde{
 u}$) llet

Molecule	Bond Order	Vibrational frequency (cm ⁻¹)	
H ₂	1	4400 ←	low mass
Li ₂	1	351	vibrational
C ₂	2	1781	frequency
N ₂	3	2358	follows
N ₂ +	2.5	2207	bond order
O ₂	2	1580	greater b.o. [
O ₂ +	2.5	1905	greater b.o. b
F ₂	1	917	greater frequency

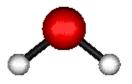
v

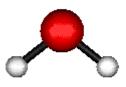
vibrational motion in molecules (H₂O)

wave number

$$\tilde{v} = \frac{1}{\lambda_{cm}} \quad \left[\tilde{v} \right] = cm^{-1}$$

$$E = hc\tilde{v}$$
 $c = 2.99792 \times 10^{10} cm sec^{-1}$





Characteristic Vibrations of H2O

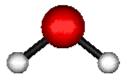
$$n_1 = 3835 \text{ cm}^{-1}$$
 $n_2 = 1648 \text{ cm}^{-1}$

$$n_2 = 1648 \text{ cm}^{-1}$$



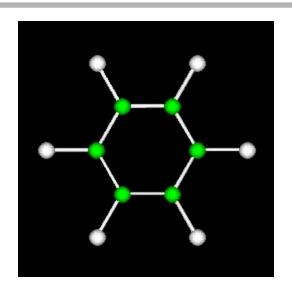


$$n_3 = 3939 \text{ cm}^{-1}$$



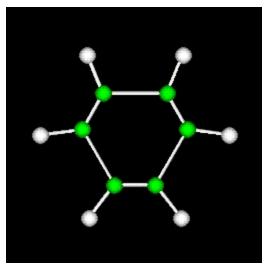


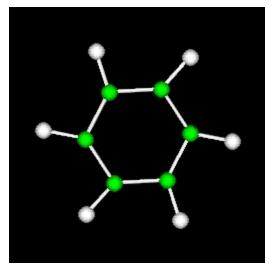
vibrational motions in molecules (benzene)

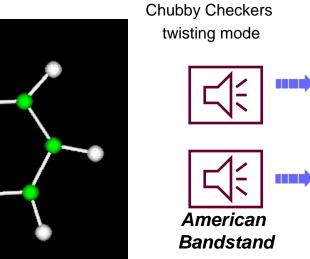


breathing (stretching) mode

asymmetric stretching mode

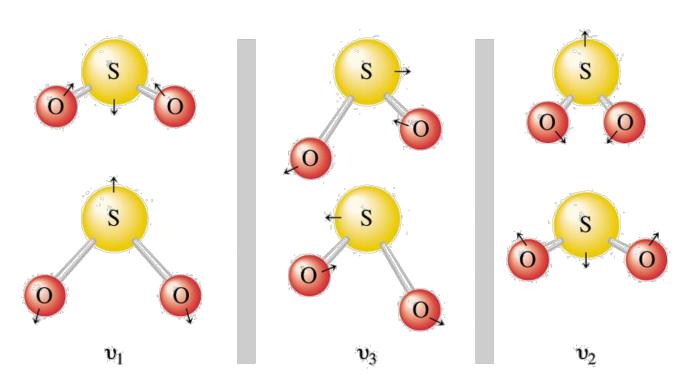








photons at infrared wavelengths excite the vibrational motion of atoms in a molecule



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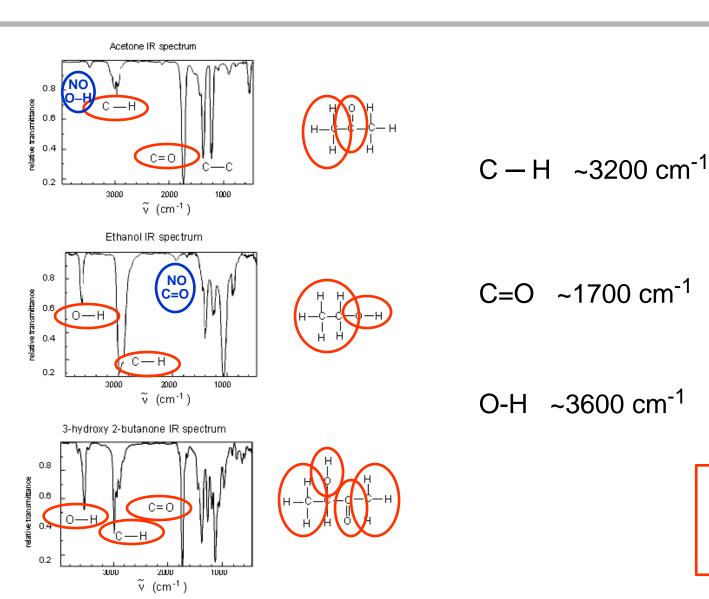
group frequencies

- different types of bonds require different energy photons for vibrational excitation
- a given bond type will have a similar absorption energy in various molecules

Bond	Characteristic Frequency (approximate) $\widetilde{v} \ (cm^{-1}) \ [E = hc\widetilde{v}]$	λ (nm)
C-C	1000-1400	10000-7140
C = C	1600	6250
$C \equiv C$	2100	4760
C – O	1100	9090
C = O	1800	5880
C — H	2800-3200	3125
O — H	3600	2770

note energy to excite bond vibration: $E_{C=C} > E_{C=C} > E_{C-C}$

IR spectra



spectroscopic excitations

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30	10 ¹⁶	far uv	vacuum UV	excitation of σ electrons
300	10 ¹⁵	near uv	UV-VIS	excitation of π and non-bonding (n) electrons
400-700	4-8 × 10 ¹⁴	visible		
3000	10 ¹³	infra-red	IR	vibrational excitations (IR)
3 × 10 ⁶	10 ¹¹	microwave	microwave ESR	rotations of molecules and flipping unpaired electron spins in external magnetic field

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3 × 10 ⁶	10 ¹¹	microwave	microwave ESR	rotations of molecules and flipping unpaired electron spins in external magnetic field
3 × 10 ⁹	10 ⁸	radiowave	NMR (MRI)	flipping of nuclear spins in an external magnetic field

NMR (MRI) spectrometers







protons (hydrogen nuclei), like electrons, behave as if they were tiny magnets

electron ~650 x stronger magnet

*The nucleus does not literally spin, although we often picture it that way. Nuclear "spin," like electron "spin," is simply a name for a property called spin angular momentum.

High energy

Low energy

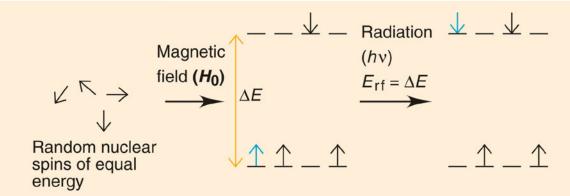
 in an external magnetic field, spin up ↑ and spin down ↓ will have different energies

FIGURE 14.62

Representations of the two spin states of the proton interacting with a large, steady mag-

 in NMR spectroscopy, photons in the radiowave region have the correct energy to cause a hydrogen nucleus to flip its spin

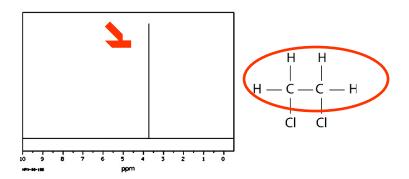
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identifying equivalent and non-equivalent protons

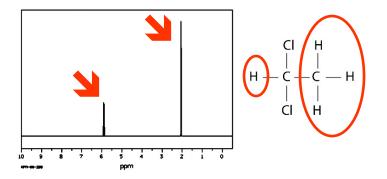
to 'flip' hydrogen atoms (nuclei) in different chemical environments requires slightly different energies (chemical shift)





(all H equivalent, H₂CIC– 1 peak one 'chemical shift')

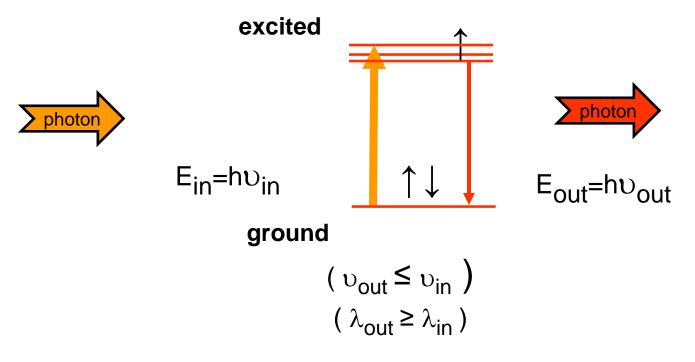
1.1 dichloroethane HNMR



(two H environments, HCl₂C- and -CH₃ 2 peaks two 'chemical shifts') not responsible for spin-spin coupling (pp. 703-704)

will get plenty in o-chem

• fluorescence- emission of radiation (almost) directly from the excited state

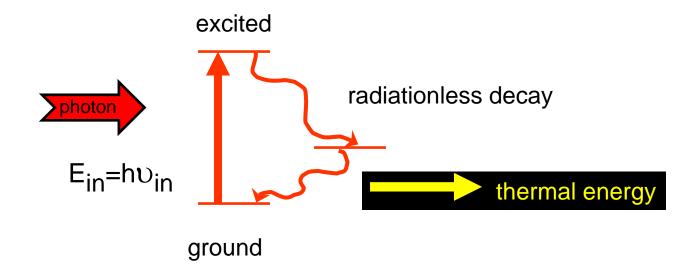


time $\sim 10^{-12}$ to 10^{-9} sec

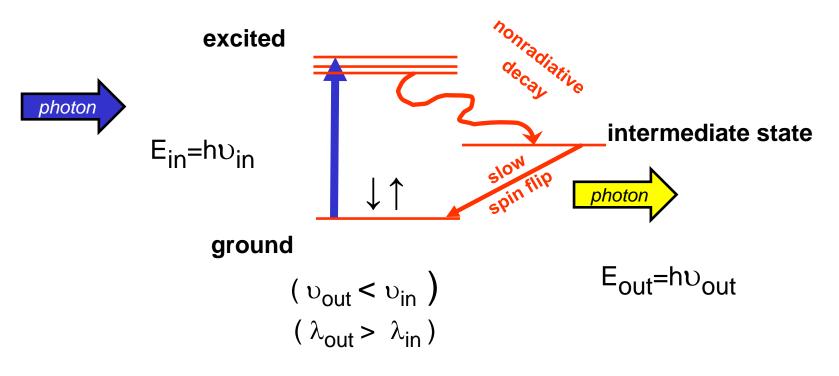
(fluorescence stops soon after exciting light is turned off)

vocabulary: radiationless decay (nonradiative decay)

 radiationless decay- transition from a higher to a lower energy state with a loss of energy in the form of heat rather than emission of a photon

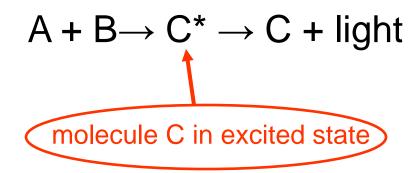


phosphorescence- 'slow' return to ground state by emission of photon from intermediate state



time $\sim 10^{-3}$ to 10 sec and longer (phosphorescence continues after exciting light is turned off)

chemiluminescence- light given off when chemical reaction leaves products in excited states and then the product fluoresces









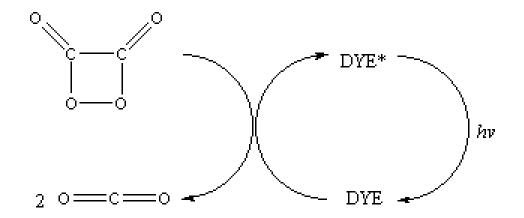


luciferin + luciferase + ATP → luciferyl adenylate-luciferase + pyrophosphate

luciferyl adenylate-luciferase + O2 → oxyluciferin + luciferase + AMP + light

chemiluminescence: light sticks



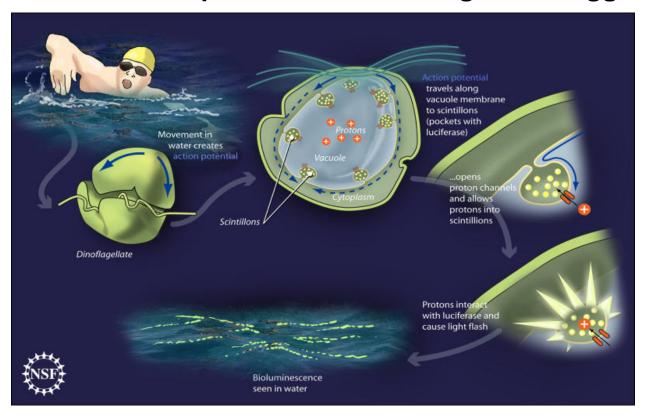


The intermediate decomposes into two carbon dioxide molecules

The fluorescer is excited; light is given off as the excited state returns to the ground state

http://www.sas.upenn.edu/~mtc/Lightstick.html

Bioluminescence: Explanation for Glowing Seas Suggested



According to the study, here is how the light-generating process in dinoflagellates may work: As dinoflagellates float, mechanical stimulation generated by the movement of surrounding water sends electrical impulses around an internal compartment within the organism, called a vacuole--which holds an abundance of protons. These electrical impulses open so-called voltage-sensitive proton channels that connect the vacuole to tiny pockets dotting the vacuole membrane, known as scintillons.

Once opened, the voltage-sensitive proton channels may funnel protons from the vacuole into the scintillons. Protons entering the scintillons then activate luciferase-a protein, which produces flashes of light, that is stored in scintillons. Flashes of light produced by resulting luciferase activation would be most visible during blooms of dinoflagellates.

end of lectures
on
spectroscopy

colored transition metal complexes- glazes



Ni(NH₃)₆Br₂

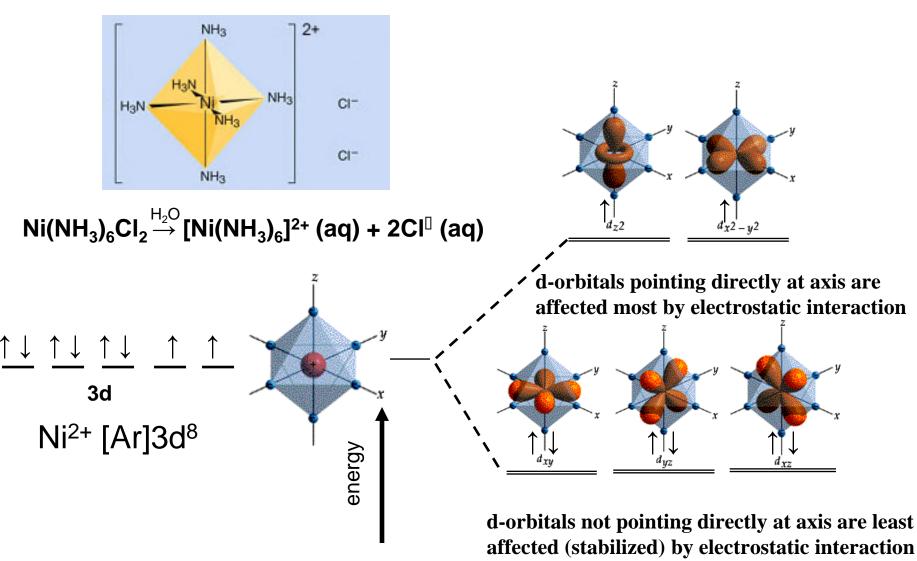


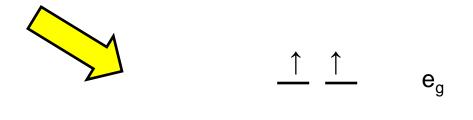
CoCl₂-6H₂O



NiSO₄-6H₂O

d-Orbitals and ligand Interaction (octahedral field)

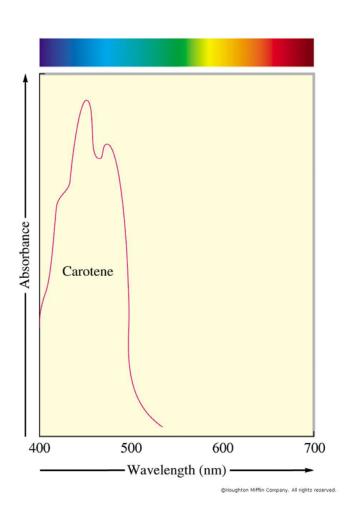


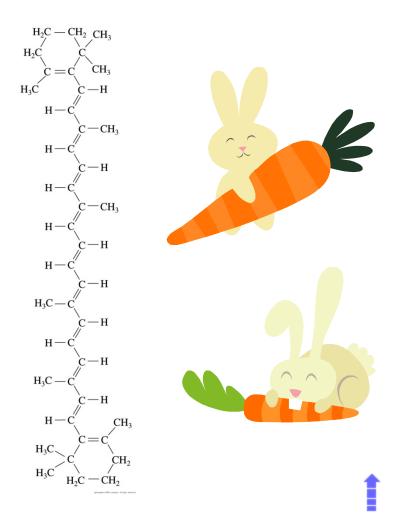


$$\frac{\uparrow \downarrow}{} \stackrel{\uparrow \downarrow}{} \stackrel{\downarrow}{} \stackrel{\downarrow}{} \stackrel{\downarrow}{} t_{2g}$$
[Ni(NH₃)₆]²⁺

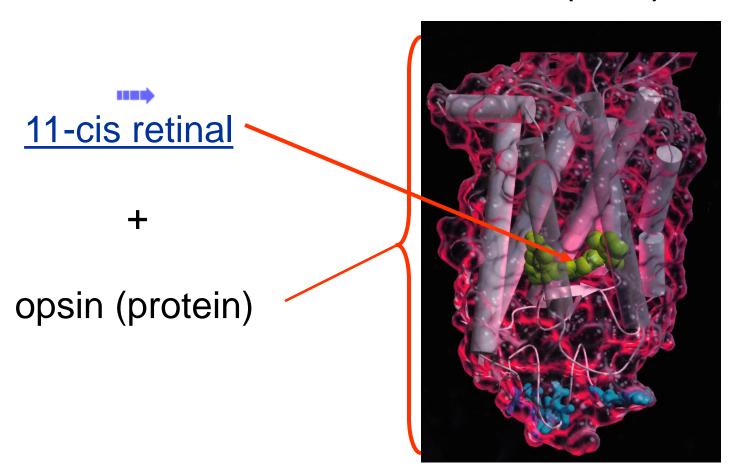
- 3d orbitals all have same energy in Ni²⁺ (g)
- presence of 6NH₃ cause "splitting of the energies of the 3d-orbitals into two levels in [Ni(NH₃)₆]²⁺
- visible light causes electronic transitions between the two levels resulting in 'colored' transition metal complexes

ß- carotene; conjugated double bonds (figure 14.56, 14.57)

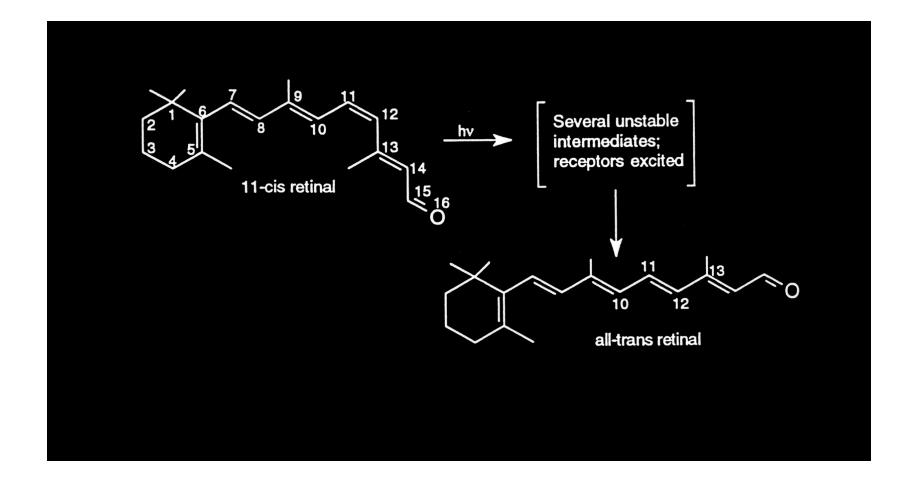


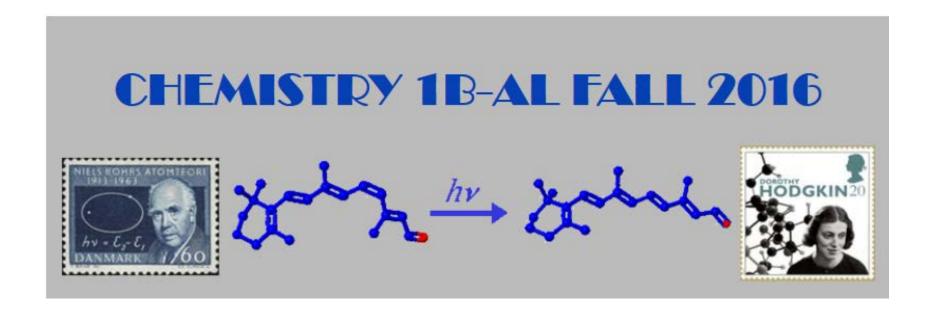


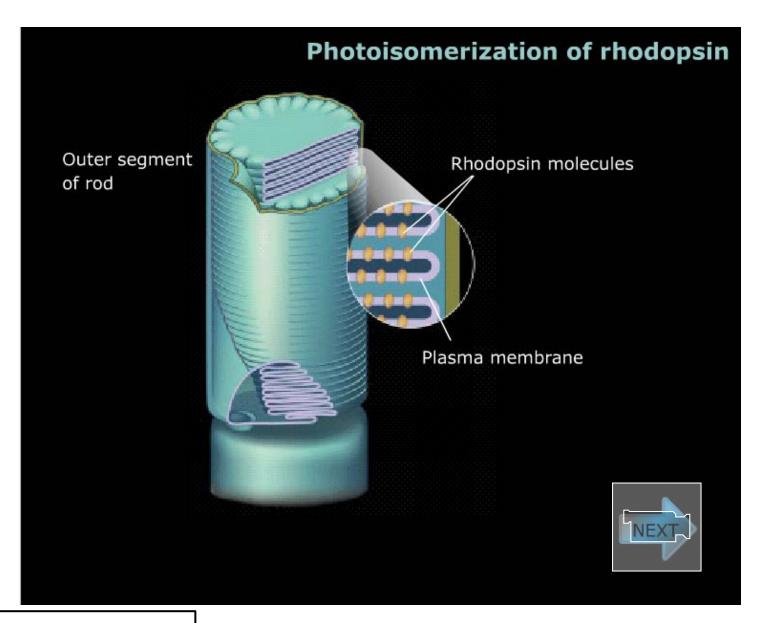
intradiscal (lumen)



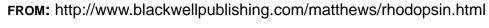
interdiscal (cytoplasmic)



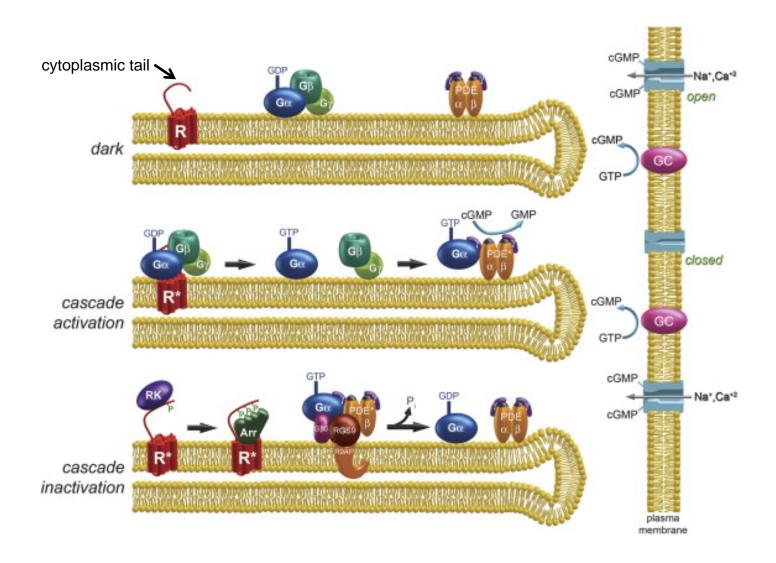








signal amplification in visual excitation cascade



amplification

